

NOTE ON UMBILICAL HYPERSURFACES WITH UNIT VECTOR FIELDS OF A REAL SPACE FORM

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Introduction.

Recently, Lawson [2] has studied a hypersurface in a real space form of constant mean curvature which has parallel Ricci tensor. With use of these results, Mogi and Nakagawa [5] have given a classification of hypersurface in a real space form with parallel Ricci tensor or the Cartan's condition about Ricci tensor.

In the present paper, we consider umbilical hypersurface M with unit vector fields in a real space form $\bar{M}(c)$, that is, there exist mutually orthogonal unit vector fields U and V such that the second fundamental tensor H of M with induced Riemannian metric tensor g has the form

$$H = \alpha I + \beta(u \otimes U + v \otimes V),$$

$$g(U, X) = u(X), \quad g(V, X) = v(X)$$

for any vector field X , α and β being functions on M .

First of all we shall prepare some local properties about a hypersurface of a real space form. In the last section 2, we prove some lemmas on an umbilical hypersurface with unit vector fields, and give classifications of the space.

§1. Certain hypersurfaces of a real space form.

Let $\bar{M}(c)$ be an $(n+1)$ -dimensional real space form covered by a system of coordinate neighborhoods $\{\bar{U}; y^a\}$, where here and in this section the indices $\lambda, \mu, \nu, \kappa, \dots$ run over the range $\{1, 2, 3, \dots, n+1\}$, that is, the curvature tensor of $\bar{M}(c)$ has the form

$$(1.1) \quad K_{\nu\mu\lambda\kappa} = c(g_{\lambda\mu}g_{\nu\kappa} - g_{\nu\lambda}g_{\mu\kappa}),$$

c being constant, where $g_{\lambda\mu}$ are components of Riemannian metric tensor of $\bar{M}(c)$.

Let M be an n -dimensional hypersurface which is covered by a system of coordinate neighborhoods $\{U; x^h\}$, where here and in the sequel the indices h, i, j, \dots run over the range $\{1, 2, 3, \dots, n\}$, and which is differentially immersed in $\bar{M}(c)$ by $X: M \rightarrow \bar{M}$, i. e., $y^\kappa = y^\kappa(x^h)$.

We put $B_i^\kappa = \partial y^\kappa / \partial x^i$, $\partial_i = \partial / \partial x^i$, then components g_{ji} of the induced metric tensor of M are given by $g_{ji} = g_{\lambda\mu} B_j^\lambda B_i^\mu$. B_i^κ are, for each i , local vector fields of $\bar{M}(c)$ tangent to M and the vectors B_i^κ are linearly independent in each coordinate neighborhood. B_i^κ is, for each κ , a local 1-form of M .

We choose a unit vector C^κ of \bar{M} normal to M in such a way that $n+1$ vectors B_i^κ , C^κ give the positive orientation of \bar{M} .

We denote $\{j^h{}_i\}$ and ∇_i by the Christoffel symbols formed with Riemannian metric g_{ji} and the operator of covariant differentiation with respect to $\{j^h{}_i\}$ respectively. Then the equations of Gauss and Weingarten are respectively

$$(1.2) \quad \nabla_j B_i^\kappa = \partial_j B_i^\kappa + \{ \mu^\kappa \lambda \} B_j^\mu B_i^\lambda - B_i^\kappa \{ j^h{}_i \} = h_{ji} C^\kappa,$$

$$(1.3) \quad \nabla_j C^\kappa = \partial_j C^\kappa + \{ \mu^\kappa \lambda \} B_j^\mu C^\lambda = -h_j^t B_t^\kappa,$$

where h_{ji} are the components of second fundamental tensor with respect to the normal C^κ , h_j^h defined by $h_j^h = h_{jt} g^{th}$ and $(g^{ji}) = (g_{ji})^{-1}$.

In the sequel, we need the structure equations of the hypersurface M , that is, the following equations of Gauss

$$(1.4) \quad K_{kjih} = c(g_{kh}g_{ji} - g_{jh}g_{ki}) + h_{kh}h_{ji} - h_{jh}h_{ki},$$

where K_{kjih} are covariant components of the curvature tensor of M , and equations of Codazzi,

$$(1.5) \quad \nabla_k h_{ji} - \nabla_j h_{ki} = 0.$$

From equations (1.4) of Gauss, we have the relationships

$$(1.6) \quad K_{ji} = (n-1)c g_{ji} + (h_t^t) h_{ji} - h_{jt} h_i^t$$

and hence

$$(1.7) \quad K = n(n-1)c + (h_t^t)^2 - h_{jt} h^{ji},$$

where K_{ji} and K are respectively the components of Ricci tensor and the curvature scalar of M .

A hypersurface M of dimension n is said to be an umbilical form with unit

vector fields, if there exist on M , two mutually orthogonal unit vector fields u^h and v^h such that

$$(1.8) \quad h_{ji} = \alpha g_{ji} + \beta(u_j u_i + v_j v_i)$$

for some functions α and β .

From the relation above, we find

$$(1.9) \quad h_i^i = n\alpha + 2\beta,$$

$$(1.10) \quad h_{ji} u^j = (\alpha + \beta) u_i, \quad h_{ji} v^j = (\alpha + \beta) v_i,$$

$$(1.11) \quad h_{ji} h^{ji} = n\alpha^2 + 4\alpha\beta + 2\beta^2$$

because u^h and v^h are unit orthogonal. Thus the second fundamental tensor (h_j^h) has at most two eigenvalues α and $\alpha + \beta$ of multiplications $n-2$ and 2 respectively.

If we substitute (1.10) and (1.11) into (1.7), we get

$$(1.12) \quad K = n(n-1)(\alpha + \beta)^2 + 4(n-1)\alpha\beta + 2\beta^2.$$

§2. Umbilical hypersurface with unit vector fields.

Throughout this paper we consider the hypersurface M of dimension $n > 3$ is an umbilical form with unit vector fields.

LEMMA 2.1. *Let M be an umbilical form with unit vector fields of dimension $n > 3$ such that the curvature scalar K is constant. Then α and β are constants on M .*

Proof. Differentiating (1.8) covariantly along M , we have

$$(2.1) \quad \nabla_k h_{ji} = \alpha_k g_{ji} + \beta_k (u_j u_i + v_j v_i) + \beta \{ (\nabla_k u_j) u_i + (\nabla_k u_i) u_j + (\nabla_k v_j) v_i + (\nabla_k v_i) v_j \},$$

from which, taking skew-symmetric parts with respect to k and j and using (1.5),

$$(2.2) \quad \begin{aligned} & \alpha_k g_{ji} - \alpha_j g_{ki} + \beta_k (u_j u_i + v_j v_i) - \beta_j (u_k u_i + v_k v_i) \\ & + \beta \{ (\nabla_k u_j - \nabla_j u_k) u_i + (\nabla_k u_i) u_j - (\nabla_j u_i) u_k \\ & + (\nabla_k v_j - \nabla_j v_k) v_i + (\nabla_k v_i) v_j - (\nabla_j v_i) v_k \} = 0, \end{aligned}$$

where $\nabla_k \alpha$ is denoted by α_k . If we transvect (2.2) with $u^i v^j v^k$ and $u^i v^j u^k$, we have respectively

$$(2.3) \quad u^t \alpha_t + u^t \beta_t = 0,$$

$$(2.4) \quad v^t \alpha_t + v^t \beta_t = 0$$

because u^h and v^h are unit orthogonal.

Differentiating (1.12) covariantly, we find

$$(2.5) \quad (n-1)(n\alpha + 2\beta)\alpha_j + 2\{(n-1)\alpha + \beta\}\beta_j = 0.$$

by virtue of $K = \text{constant}$, from which, transvecting u^j and using (2.3),

$$(2.6) \quad (n-2)\{(n-1)\alpha + 2\beta\}(\alpha, u^t) = 0.$$

If $\alpha, u^t \neq 0$, then $(n-1)\alpha + 2\beta = 0$ which implies $(n-3)\nabla_j(\alpha^2) = 0$. This contradicts $\alpha, u^t \neq 0$. Consequently we have

$$(2.7) \quad \alpha, u^t = 0, \quad \beta, u^t = 0.$$

In the same way we also have from (2.4)

$$(2.8) \quad \alpha, v^t = 0, \quad \beta, v^t = 0.$$

Next, transvecting (2.2) with g^{ji} and taking account of (2.7) and (2.8), we obtain

$$(2.9) \quad (n-1)\alpha_k + 2\beta_k \\ = \beta\{u^t \nabla_t u_k + v^t \nabla_t v_k + (\nabla_t u^t)u_k + (\nabla_t v^t)v_k\}.$$

On the other hand, if we transvect (2.2) with $u^j u^i$ and $v^j v^i$, we get respectively

$$(2.10) \quad \alpha_k + \beta_k + \beta\{-u^t \nabla_t u_k - (u^s u^t \nabla_t v_s)v_k\} = 0,$$

$$(2.11) \quad \alpha_k + \beta_k + \beta\{-v^t \nabla_t v_k - (v^s v^t \nabla_t u_s)u_k\} = 0.$$

Combining (2.9), (2.10) and (2.11), we conclude

$$(2.12) \quad (n-3)\alpha_k = -\beta\{(v^s v^t \nabla_t u_s - \nabla_t u^t)u_k + (u^s u^t \nabla_t v_s - \nabla_t v^t)v_k\},$$

which implies that $v^s v^t \nabla_t u_s = \nabla_t u^t$, $u^s u^t \nabla_t v_s = \nabla_t v^t$ because of (2.7) and (2.8). Thus (2.12) means α is constant for $n > 3$ and hence β is also by virtue of (2.5). Therefore, Lemma 2.1 is proved.

LEMMA 2.2. *Under the same assumptions as those stated in Lemma 2.1 we have $\nabla_k h_{ji} = 0$ and consequently $\nabla_k K_{ji} = 0$.*

Proof. α and β being constants because of Lemma 2.1, we see from (2.2)

that $\beta=0$ or

$$(2.13) \quad \begin{aligned} &(\nabla_k u_j - \nabla_j u_k) u_i + (\nabla_k u_i) u_j - (\nabla_j u_i) u_k \\ &+ (\nabla_k v_j - \nabla_j v_k) v_i + (\nabla_k v_i) v_j - (\nabla_j v_i) v_k = 0. \end{aligned}$$

If $\beta=0$, then M is totally umbilical by virtue of (1.8) and hence $\nabla_k h_{ji}=0$. Thus we may only consider $\beta \neq 0$.

Transvecting (2.13) with u^i and v^i , we obtain respectively

$$(2.14) \quad \nabla_k u_j - \nabla_j u_k = A_j v_k - A_k v_j,$$

$$(2.15) \quad \nabla_k v_j - \nabla_j v_k = A_k u_j - A_j u_k,$$

where $A_j = u^t \nabla_j v_t$.

From (2.14) and (2.15) we have

$$(2.16) \quad v^t \nabla_t u_j = -(v^t A_t) v_j, \quad u^t \nabla_t u_j = -(u^t A_t) v_j,$$

$$(2.17) \quad u^t \nabla_t v_j = (u^t A_t) u_j, \quad v^t \nabla_t v_j = (v^t A_t) u_j.$$

Substituting (2.14) and (2.15) into (2.13), we obtain

$$(2.18) \quad \begin{aligned} &A_k(u_j v_i - v_j u_i) + A_j(v_k u_i - u_k v_i) \\ &+ u_j \nabla_k u_i - u_k \nabla_j u_i + v_j \nabla_k v_i - v_k \nabla_j v_i = 0. \end{aligned}$$

Transvecting (2.18) with u^j, v^j and taking account of (2.16) and (2.17), we find respectively

$$(2.19) \quad \nabla_k u_i = -A_k v_i,$$

$$(2.20) \quad \nabla_k v_i = A_k u_i.$$

Thus (2.1) implies $\nabla_k h_{ji}=0$ because of (2.19), (2.20) and Lemma 2.1. Thus (1.6) proves the last assertion of the lemma.

LEMMA 2.3. *Under the same assumptions as those stated in Lemma 2.1, we have*

$$(2.21) \quad \alpha(\alpha + \beta) + c = 0.$$

Proof. Differentiating (2.19) covariantly and using (2.20), we get

$$(2.22) \quad \nabla_k \nabla_j u_i = -(\nabla_k A_j) v_i - A_j A_k u_i,$$

from which, taking skew-symmetric parts with respect to k and j and making use of the Ricci identity,

$$(2.23) \quad -K_{kji}u^k = (\nabla_j A_k - \nabla_k A_j)v_i,$$

or, using (1.4) and (1.10),

$$(2.24) \quad \begin{aligned} & (\nabla_k A_j - \nabla_j A_k)v_i \\ &= c(u_k g_{ji} - u_j g_{ki}) + (\alpha + \beta)(u_k h_{ji} - u_j h_{ki}). \end{aligned}$$

Transvecting (2.24) with $u^k v^j v^i$, $g^{ji} u^k$ and using (1.9) and (1.10), we have respectively

$$\begin{aligned} u^i v^i \nabla_i A_s - u^i v^i \nabla_s A_i &= c + (\alpha + \beta)^2, \\ u^i v^i \nabla_i A_s - u^i v^i \nabla_s A_i &= (n-1)c + (\alpha + \beta) \{ (n\alpha + 2\beta) - (\alpha + \beta) \}. \end{aligned}$$

The last two relations imply (2.21). This completes the proof of the lemma.

In the case where ambient space \bar{M} is Euclidean, from (1.12) and (2.21), we have

$$(2.25) \quad K = (n-1)(n-4)\alpha^2 + 2\beta^2 \geq 0.$$

If the curvature scalar K is positive, by completeness, M is congruent to $S^2(r) \times E^{n-2}$ or $S^{n-2}(r) \times E^2$, and if $K=0$, M is cylindrical because the Ricci tensor is parallel (cf. [3], [5]).

Thus we have proved

THEOREM 2.4 *Let M be a complete and connected umbilical hypersurface with unit vector fields defined by (1.8) such that $\dim M > 3$ and the curvature scalar K is constant. Then M is congruent to $S^2(r) \times E^{n-2}$ or $S^{n-2}(r) \times E^2$ if the scalar curvature $K > 0$, and M is a cylinder if the scalar curvature $K = 0$.*

Now, we suppose that the real space form $\bar{M}(c)$ is of constant curvature $c \neq 0$ and the hypersurface M has the constant scalar curvature \bar{K} and $n > 3$. Then by means of Lemma 2.1, 2.2 and 2.3, we have two cases: (1) M has exactly two distinct constant principal curvatures, say α and $\alpha + \beta$ of multiplicities $n-2$ and 2 respectively, such that $c + \alpha(\alpha + \beta) = 0$, and (2) M is totally umbilic but not totally geodesic.

For the first case, we use Lemma 2.3. Then, from the straightforward argument used by Lawson [2], we obtain the following conclusion:

If $c > 0$, then M is isometric to $S^2(c_1) \times S^{n-2}(c_2)$, and if $c < 0$, then M is isometric to $S^2(c_1) \times H^{n-2}(c_2)$, $S^r(a)$ being a sphere with curvature c and $H^r(a)$ a hyperbolic space with curvature a .

For the second case, M is totally umbilic but not totally geodesic. If $c > 0$, then M is isometric to a sphere S^n , and if $c < 0$, then M is a sphere S^n , a hyperbolic space H^n whose curvature is different from c , or a flat hypersurface F^n .

Thus, summing up the results obtained above, we have proved

THEOREM 2.5. *Let \bar{M} be an $(n+1)$ -dimensional and simply connected real space form with curvature $c \neq 0$ and let $M (n > 3)$ be a complete and connected umbilical hypersurface with unit vector fields defined in (1.8) such that the curvature scalar K is constant. Then the following statements are true:*

(1) *If $c > 0$, then M is isometric to the great sphere, the small sphere or $S^2(c_1) \times S^{n-2}(c_2)$, where $1/c_1 + 1/c_2 = 1/c$.*

(2) *If $c < 0$, then M is isometric to S^n , H^n , F^n or $S^2(c_1) \times H^{n-2}(c_2)$, $S^{n-2}(c_1) \times H^2(c_2)$, where $1/c_1 + 1/c_2 = 1/c$.*

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