

NASA's coronagraph-monitoring programs. Ground-based data show global-scale outflows of Na in contrast to small-scale data usually obtained near the orbits of spacecraft.

## 항성 및 항성계 / 고천문학

### [구 SH-01] Evolutionary properties of red supergiants with MESA

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We investigate the evolutionary properties of red supergiant stars (RSGs), using stellar evolution model of Modules for Experiments in Stellar Astrophysics (MESA). In this study, we calculate models with mass range of 9-39M<sub>sun</sub> and several different convection parameters (e.g. mixing length, overshooting, and semiconvection) at SMC, LMC, Milky Way, and M31 metallicities. We compare the calculated evolutionary tracks with observed RSGs in SMC, LMC, Milky Way and M31, and discuss appropriate input physical parameters in model calculation. We find that a larger mixing length parameter is necessary for M31 metallicity to fit the positions of RSGs in H-R diagram, compared to lower metallicity environments. Theoretically predicted numbers of yellow supergiant stars (YSGs) are also compared with the observed population. We find that Ledoux models with semiconvection can better explain the number of YSGs. Finally, we investigate the final radius, final star mass, and final hydrogen envelope mass of RSGs and discussed the their properties as type II-P supernova progenitors.

### [구 SH-02] The CN-CH positive correlation in the globular cluster NGC 5286

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We performed low-resolution spectroscopy for the red giant stars in the Galactic globular cluster (GC) NGC 5286, which is known to show intrinsic heavy element variations. We found that these stars are clearly divided into two subpopulations

by CN index. These two subpopulations also show significant differences in the HK' and CH indices, where the CN-strong stars are more enhanced in both indices. From the comparison with high-resolution spectroscopic data of Marino et al. (2015), we found that the CN- and HK'-strong stars are also increased in the abundances of s-process elements and Fe. It appears that, therefore, these stars are later generation stars probably affected by supernova enrichment. In addition, NGC 5286 shows the CN-CH positive correlation among the whole sample, which is only discovered in the GCs with heavy element variations such as M22 and NGC 6273. Therefore, these results strengthen our previous suggestion that the CN-CH positive correlation may be associated with the heavy element variations in the GCs.

### [구 SH-03] Origin of Low-mass Hypervelocity Stars in the Galactic Disk

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We present the analysis of chemical abundances and kinematics for six hypervelocity star (HVS) candidates. These objects are G/K-type low-mass stars in the Galactic disk, while other HVSs previously found are B-type high-mass objects in the Galactic halo. The stellar orbits and kinematics of our HVS candidates suggest that they do not originate in the Galactic center or in an accretion event, indicative of yet-unknown mechanisms that produce kinematically-extreme disk stars. In order to study in detail their origin, we obtained medium-resolution (R~6000) spectra of these stars and derived abundances of several chemical elements (Mg, Ca, Si, Ti, Cr, Fe, and Ni). From the comparison of the chemical abundances with the Galactic stellar components (disk, bulge, halo, and dwarf galaxies) and the kinematic properties of our HVSs, we conclude that two of them are likely ejected from the Galactic disk, one originated from the Galactic center as for the young B-type HVSs, and the other one might be ejected from either the Galactic disk or other regions.

### [구 SH-04] ANALYSIS OF LONG PERIOD RADIAL VELOCITY VARIATIONS FOR HD 18438 AND HD 158996

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We investigate the long-period radial velocity (RV) variations for M giant HD 18438 and

K giant HD 158996 using the high-resolution Bohyunsan Observatory Echelle Spectrograph at the 1.8m telescope of Bohyunsan Optical Astronomy Observatory in Korea. These two target stars are important because HD 18438 is the largest star and HD 158996 is the brightest star for exoplanetary system candidate so we can understand how evolved stars affect planets by researching these stars. We calculated precise RV measurements of 38 and 24 spectra from November 2010 to January 2017 and June 2010 to January 2017, respectively. We derived the RV variation period for 719.0 days of HD 18438, 775.6 days for HD 158996. We conclude that the RV variation of HD 158996 is caused by planetary companion which has the mass of 14.7 MJup, semi-major axis of 2.2 AU, and eccentricity of 0.27 assuming the stellar mass of 2.34 M $\odot$ . On the other hand, the origin of RV variation of HD 18438 with period of 719.0 days is still uncertain. It might be caused by stellar chromospheric activity or planetary companion, so more observations and tests are required.

### [구 SH-05] Black Hole Binaries Dynamically Formed in Globular Clusters

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We investigate properties of black hole (BH) binaries formed in globular clusters, by using direct N-body simulations. Comparing with previous studies which usually considered single BH masses, our models consist of two-component BH masses, or continuous BH mass function with single mass ordinary cluster stars. During the early stage of dynamical evolution, initially distributed BHs are move to the cluster center by dynamical friction, then BH-BH binaries start to be

formed, and eventually be ejected from the cluster due to three body interaction. Finally we find the formation efficiency of high mass BHs are always larger than that of lower mass BHs, implying that a BH mass spectrum expected from GW observation should be biased to high mass. In addition, mass ratios of BHs in binaries prefer similar masses (ratio~1), while the most extreme case is less than 3. Expected merger rate from our models is about 7 BH-BH mergers per Mpc<sup>3</sup> per yr.

### [구 SH-06] The Constellation Maps in the Flags of Barracks in GANGJIN

(강진 병영 영기(令旗)에 그려진 별자리)

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전라남도 강진에서 발견된 영기(令旗)라는 책에는 병영에서 사용된 별자리 깃발에 관한 기록이 남아있다. 영기는 군중에서 군령(軍令)을 전달하기 위해 사용한 것으로 고종대에 이르러 깃발에 28수(宿) 별자리를 처음으로 사용한 것으로 알려져 있다. 승정원일기와 일성록에 의하면 군영에서 사용한 28수 별자리 깃발은 1874년 중앙관 진무사(鎭撫使)의 수장이었던 김선필(金善弼)이 처음 만들어 사용한 것으로 기록되어 있다. 본 발표에서는 국내에 처음 보고된 28수(宿)가 그려진 영기를 소개하고 영기의 별그림을 한국과 중국의 전통 성도와 비교한 내용을 발표하고자 한다. 영기에는 28수 별자리 외에도 28 동물과 기하학적 문양이 그려져 있는데 이에 대해서도 간단히 소개하고자 한다. 영기 별그림은 실제 성도와 비교해 많은 오류가 확인되는데 이를 교정하여 새롭게 도안한 별자리 깃발도 함께 소개하고자 한다.

**KMTNet**

### [구 KMT-01] Operation and System Upgrade of KMTNet

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We report the operational highlights of KMTNet in the point of observing rate, image pre-processing and data reduction, observing run for each science program, and scientific publications performed in 2016. Major system upgrade has been conducted in the CCD camera and the wide field telescope optics: the post amp and readout electronics of the 18k Mosaic CCD