Skin Effect of Rotating Magnetic Fields in Liquid Bridge

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A rotating magnetic field (RMF) $\Phi_1-\Phi_2$ model was developed in consideration of the skin effect. The rotating magnetic field’s induced three-dimensional flow was simulated numerically, and the influence of the skin effect was investigated. The rotating magnetic field drives the rotating convection in the azimuthal direction, and a secondary convection appears in the radial-meridional direction. The results indicate that ignoring the skin effect results in a smaller azimuthal velocity component and larger radial and axial velocity components, and that the deviation becomes more obvious with the larger dimensionless shielding parameter $K$.

Keywords: Rotating magnetic Fields, Computer simulation, Fluid flows, Skin effect, $\Phi_1-\Phi_2$ model

1. Introduction

With the rapid development of information technology, there has been a growing demand for large and high-quality semiconductor single crystals. During crystal growth, the heat and mass transfer of melt play important roles in determining crystal quality. In particular, for large crystals, convection control becomes significant to the achievement of high quality. Nowadays, the magnetic field, either static \cite{1, 2} or rotating, is essential to the control of melt convection. However, a rotating magnetic field, relative to a static one, requires much less energy, and therefore attracts increasing attention in the research field \cite{3-8}.

In a rotating magnetic field, the skin effect leads to a non-uniform magnetic field in melt and depends on the shielding parameter, $K = \sigma \mu_0 \omega R^2$ (where $\sigma$ is the electrical conductivity, $\mu$ is the magnetic permeability, $\omega$ is the rotating angular frequency, and $R$ is the radius of melt). The shielding parameter, $K$, is usually used to characterize the interaction between the magnetic field and the electrically conducting melt. The condition $K \ll 1$ means that the magnetic field distribution is almost completely unaffected by the conducting melt; in other words, the rotating magnetic field is taken to penetrate into the melt volume without any change, under which condition the skin effect is negligible. In most of the research on the rotating magnetic field for convection control in crystal growth, the dimensionless shielding parameter $K$ is assumed to be $K \ll 1$ \cite{3-8}. Nevertheless, in more typical cases, $K$ is expected to be larger than 1, which indicates that the magnetic field is reduced when penetrating into the melt, because the magnetic field lines are expelled due to the high conductivity of the melt and the large angular frequency of the magnetic field. In the growth of large-size crystals, the dimensionless shielding parameter $K$ can be large, for example, 8.88 for a rotating magnetic field of 50 Hz frequency in a silicon melt of 15 cm radius. In this case, the skin effect of the rotating magnetic field should be studied.

In this case, the skin effect in the rotating magnetic field has been investigated in \cite{9-14}. E. Dahlberg \cite{9} reported the expression of the Lorentz force with the skin effect by ignoring the effects of the convection on the electric current and the magnetic field. Volz and Mazuruk \cite{11} theoretically studied the Lorentz force in melt within a
finite-length cylinder under a rotating magnetic field, and obtained an approximate solution of the azimuthal Lorentz force by ignoring the interaction between the field and the melt convection. Spitzer et al. [12, 13] presented approximate expressions of Lorentz force in (the radial and azimuthal directions) in Czochralski crystal growth while the effects of convection on the electric current and magnetic field were ignored. In all of the above investigations, the infinite model was adopted to calculate the distribution of the magnetic field. Moreover, in order to derive the Lorentz force, the following assumptions were taken into account: the height of the melt is infinite; the contribution of melt convection to the Lorentz force is negligible, and only the azimuthal Lorentz force is derived. However, Yao et al. [8] stated that the over-simplified RMF infinite model might lead to large deviations in convection structure and temperature distribution, because the components of the Lorentz force in both the radial and axial directions can affect the melt flow.

In the present study, in calculating the distribution of the magnetic field, the assumption of an infinite-length inductor of cylindrical symmetry was taken, as in the preceding reports, but in subsequently deriving the Lorentz force, all of the other above-noted assumptions were abandoned. The rotating magnetic field $\Phi_1-\Phi_2$ model, including all of the components of the Lorentz force, was derived, and the interaction between the melt convection and the magnetic field was considered in the developed rotating magnetic field $\Phi_1-\Phi_2$ model. Then, the skin effect of the rotating magnetic field was investigated via a three-dimensional flow driven by that field.

2. Physical and Mathematical Model

The geometrical model, as exhibited in Fig. 1, is a cylindrical liquid bridge suspended between two disks. The rotating magnetic field with strength $B_0$ rotates with angular frequency $\omega$ in the azimuthal direction, and the convection is driven by the rotating magnetic field due to the Lorentz force. The height and radius of the liquid bridge are $H$ and $R$, respectively. Basically, both the magnetic strength and the gradient of the rotating magnetic field in the axial direction ($z$ direction) are assumed to be zero.

The applied external rotating magnetic field in the Cartesian coordinate system is

$$\vec{B} = B_0 \left[ (\cos \omega t \cdot \hat{e}_x + \sin \omega t \cdot \hat{e}_y) \right],$$

(1)

where $\hat{e}_x, \hat{e}_y$ are unit vectors in the $x$ and $y$ directions, respectively.

The rotating magnetic distribution, as express by equation (1), does not depend on $(x, y)$; when the skin effect is considered, as in the present study, the rotating magnetic field is reduced into the melt, the magnetic field depends on $(x, y)$, and its distribution is given by the solution of the following advection-diffusion equation:

$$\frac{\partial \vec{B}}{\partial t} = \nabla \times (\vec{U} \times \vec{B}) + \frac{1}{\mu \sigma} \nabla^2 \vec{B},$$

where $\vec{U}$ is the fluid velocity and $\mu$ is the melt permeability. The first term in the right represents the magnetic flux density varies with fluid flow, and this term is ignored as in [15]. Thus, the above advection-diffusion equation is simplified as

$$\frac{\partial \vec{B}}{\partial t} = \frac{1}{\mu \sigma} \nabla^2 \vec{B}.$$  

(2)

In addition, the magnetic field satisfies

$$\nabla \cdot \vec{B} = 0.$$  

(3)

To calculate the distribution of magnetic field in melt, the cylindrical coordinate system is adopted. The following complex variables are introduced, where the real part corresponds to the physical quantities. As $r \to \infty$, the imposed magnetic field is written as

$$\vec{B}_{r=\infty} = B_0 e^{i(\omega t-\nu)}(\hat{e}_r - i\hat{e}_\theta),$$

(4)

Fig. 1. (a) Liquid bridge model, and (b) external rotating magnetic field.
where, $t$ denotes time, $\theta$ is angular coordinate, $i$, with $i^2 = -1$, is imaginary unit, $\hat{e}_r$ and $\hat{e}_\theta$ are unit vectors in radial and azimuthal directions, respectively.

Equation (3) can be written in cylindrical coordinate system as,

$$\frac{\partial}{\partial t} (r B_r) + \frac{1}{r} \frac{\partial}{\partial \theta} (r B_\theta) = 0 \ .$$  \hspace{1cm} (5)

Combining (2) and (5) in cylindrical coordinate system, we obtain

$$\frac{\partial^2}{\partial t^2} (r B_r) + \frac{1}{r} \frac{\partial}{\partial r} \left[ \frac{1}{\sigma \mu} r \frac{\partial}{\partial r} (r B_r) + \frac{1}{r^2} \frac{\partial^2}{\partial \theta^2} (r B_r) \right] = 0 \ .$$  \hspace{1cm} (6)

From the equation (4), an approximate boundary condition at $r = R$ is determined as

$$B_r \bigg|_{r=R} = B_0 e^{i(\omega t - \theta)} \ .$$  \hspace{1cm} (7)

Based on the equation (6), $B_r$ is assumed in the form as

$$B_r = \frac{f(r)}{r} e^{i(\omega t - \theta)} \ ,$$  \hspace{1cm} (8)

where $f$ is a function of $r$. From equations (5) and (8), we derive

$$B_r = -i \frac{d}{dr} f e^{i(\omega t - \theta)} \ .$$  \hspace{1cm} (9)

Substituting equation (8) into equation (6), the equation for $f$ becomes

$$r^2 \frac{d^2 f}{dr^2} + r \frac{df}{dr} + \left[ \frac{\sigma \mu \omega r^2}{1} + 1 \right] f = 0 \ .$$  \hspace{1cm} (10)

Defined $\beta = \sqrt{\sigma \mu \omega e^{3/4}}$, the equation (10) is written as

$$r^2 \frac{d^2 f}{dr^2} + r \frac{df}{dr} + \left[ \beta^2 r^2 - 1 \right] f = 0 \ .$$  \hspace{1cm} (11)

With definition of a new independent variable $\hat{r} = \beta r$, the equation (11) becomes

$$\hat{r}^2 \frac{d^2 f}{d\hat{r}^2} + \hat{r} \frac{df}{d\hat{r}} + \left[ \hat{r}^2 - 1 \right] f = 0 \ ,$$  \hspace{1cm} (12)

it is a standard Bessel equation, and the boundary condition for $f$ is derived from (7) and (8) as $f \bigg|_{r=R} = B_0 R$.

The equation (12) is solved with Bessel functions as

$$f = \frac{R B_r J_1 \left( \frac{\sqrt{K}}{R} r \sqrt{-i} \right)}{J_1(\sqrt{K} \sqrt{-i})} \ .$$  \hspace{1cm} (13)

The physical values of $B_r$ and $B_\theta$ are the real parts of the equations (13) and (14), i.e.

$$B_r = \Re \left\{ \frac{R B_r J_1 \left( \frac{\sqrt{K}}{R} r \sqrt{-i} \right)}{J_1(\sqrt{K} \sqrt{-i})} \right\} \ ,$$  \hspace{1cm} (15)

$$B_\theta = \Re \left\{ -i e^{i(\omega t - \theta)} \frac{d}{dr} \left( \frac{R B_r J_1 \left( \frac{\sqrt{K}}{R} r \sqrt{-i} \right)}{J_1(\sqrt{K} \sqrt{-i})} \right) \right\} \ .$$  \hspace{1cm} (16)

Here, the Kelvin functions $ber_1(x)$ and $bei_1(x)$ are the real and imaginary parts, respectively, of $J_1(x \sqrt{-i})$, where $x$ is real and $J_1(x \sqrt{-i})$ is the first order Bessel function of the first kind.

In order to model the interaction of the rotating magnetic field and melt convection, the above distribution of magnetic field is introduced into $\Phi_1 - \Phi_2$ model. To facilitate calculating the Lorenz force and programming the rotating magnetic field $\Phi_1 - \Phi_2$ model, the Cartesian coordinate is adopted, and the power series expansion of the analytic solution (i.e. Eqs. (15-16)) is truncated with reserving different terms. In this study, we reserve only the first four terms in the power series expansions of the Kelvin functions as

$$ber_1(x) = \Re [J_1(x \sqrt{-i})] =$$

$$- \frac{1}{\sqrt{2}} \left[ \frac{x}{2} + \frac{(x/2)^3}{1!2!} - \frac{(x/2)^5}{2!3!} + \frac{(x/2)^7}{3!4!} \right] \ ,$$

$$bei_1(x) = \Im [J_1(x \sqrt{-i})] =$$

$$- \frac{1}{\sqrt{2}} \left[ \frac{x}{2} + \frac{(x/2)^3}{1!2!} + \frac{(x/2)^5}{2!3!} - \frac{(x/2)^7}{3!4!} \right] \ .$$

As our analysis in [16], above truncation of first four terms is applicable for $K \leq 8$. In the Cartesian coordinate system, the magnetic field with the skin effect is derived as
\[
\vec{B}_{\text{rot}}(x, y, t) = \frac{B_0}{2 + \frac{K^2}{96} + \frac{2K^5}{9216}} \left[ (\sin \omega t \cdot B_{\text{sin}} + \cos \omega t \cdot B_{\text{cos}}) \hat{e}_x \right] + (\sin \omega t \cdot B_{\text{y,sin}} + \cos \omega t \cdot B_{\text{y,cos}}) \hat{e}_y \],
\]
where \(B_{\text{sin}}, B_{\text{y,sin}}\) and \(B_{\text{cos}}\) are derived in Appendix A.

### 2.1. Rotating magnetic field \(\Phi_1 - \Phi_2\) model without the skin effect

The Lorentz force is obtained by \(\vec{f} = \vec{j} \times \vec{B}\), where the current density \(\vec{j}\) is determined by \(\vec{j} = \sigma(\vec{E} + \vec{U} \times \vec{B})\). In rotating magnetic field, the electric field \(\vec{E}\) is derived from \(\nabla \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}\) and \(\vec{B} = \nabla \times \vec{A}\) as
\[
\vec{E} = -\nabla \phi + (B_0 \sigma_0 \cos \omega t + B_0 \alpha_0 y \sin \omega t) \hat{e}_z .
\]

The electric field \(\vec{E}\), the vector potential \(\vec{A}\), the current density \(\vec{j}\) and the scalar electrical potential \(\phi\) depend on the angular frequency \(\omega\). In particular, \(\phi\) is split into two parts as in [3, 5]:
\[
\phi(x, y, z, t) = \phi(x, y, z) \sin(\omega t) + \phi(x, y, z) \cos(\omega t) .
\]

The Lorentz force includes both time-independent and time-dependent terms, the time-dependent term rotates at an angular frequency of 2\(\omega\). Since the melt flow cannot respond quickly at the angular frequency 2\(\omega\), the Lorentz force is taken as a time average of one rotation period, and therefore only the time-independent term of the Lorentz force is preserved as in [3, 6, 8]. The components of the Lorentz force are
\[
f_x = -\frac{1}{2} \sigma \omega B_0 \sigma_0 y + \frac{1}{2} \sigma \omega \frac{\partial \phi}{\partial z} - \frac{1}{2} \sigma B_0^2 u ,
\]
\[
f_y = \frac{1}{2} \sigma \omega B_0 \sigma_0 x - \frac{1}{2} \sigma \omega \frac{\partial \phi}{\partial z} + \frac{1}{2} \sigma B_0^2 v ,
\]
\[
f_z = \frac{1}{2} \sigma B_0 \left( \frac{\partial \phi}{\partial y} - \frac{\partial \phi}{\partial x} \right) - \sigma B_0^2 w .
\]

Wherein, \(u, v, w\) are components of the velocity in \(x, y, z\) directions, respectively.

Taking \(\nabla \cdot \vec{j} = 0\), the governing equation of scalar electric potential is derived as
\[
\nabla^2 \phi = B_0 \left( \frac{\partial u}{\partial z} - \frac{\partial w}{\partial x} \right) ,
\]
\[
\nabla^2 \phi_x = B_0 \left( \frac{\partial w}{\partial y} - \frac{\partial v}{\partial z} \right) .
\]

Electrically insulating boundaries are assumed to surround the liquid bridge, and the boundary condition is derived from \(\vec{j} \cdot \hat{n} = 0\) with the unit normal vector directing outside of the boundaries \(\hat{n}\).

The melt is taken as incompressible Newtonian fluid with constant viscosity \(\eta\) and density \(\rho\). Moreover, this study concentrates on the influence of rotating magnetic field, and the buoyancy and thermocapillary convections are ignored. Employing \(R\) and \(k/R\) as the scale of length and velocity, the dimensionless governing equations are derived as:
\[
\nabla \cdot \vec{U}' = 0 ,
\]
\[
\frac{1}{\text{Pr}} \left( \frac{\partial \vec{U}'}{\partial t'} + (\vec{U}' \times \nabla) \vec{U}' \right) = -\nabla P' + \nabla \phi' + \vec{f}_{\text{Lorz}}' .
\]
\[
\frac{\partial T'}{\partial t'} + (\vec{U}' \times \nabla) T' - \nabla^2 T' = 0
\]

where \(\vec{U}' = (R / \kappa) \vec{U} , \quad t' = (k / R^2) t , \quad \rho' = (R^2 / \kappa \eta) \rho , \quad \phi' = (1 / B_0 \kappa) \phi . \quad \kappa\) and \(\sigma\) are thermal diffusion coefficient and the electrical conductivity, respectively.

The Prandtl, rotation Reynolds and Taylor numbers are defined as
\[
\text{Pr} = \frac{\nu}{\kappa} , \quad \text{Re}_a = \frac{\omega R^2}{\nu} \quad \text{and} \quad \text{Ta} = \frac{\sigma B_0^2 \omega R^4}{2 \nu \eta} ,
\]
where \(\nu\) is the kinematic viscosity with \(\nu = \eta / \rho\).

Three components of the dimensionless Lorentz force in the rotating magnetic field \(\Phi_1 - \Phi_2\) model without the skin effect are
\[
f_x' = \text{Ta} \text{Pr} \left[ -y' + \frac{1}{\text{Re}_a} \text{Pr} \left( \frac{\partial \phi'}{\partial z} - u' \right) \right] ,
\]
\[
f_y' = \text{Ta} \text{Pr} \left[ x' - \frac{1}{\text{Re}_a} \text{Pr} \left( \frac{\partial \phi'}{\partial z} + v' \right) \right] ,
\]
\[
f_z' = \frac{\text{Ta}}{\text{Re}_a} \left[ \frac{\partial \phi'}{\partial y} - \frac{\partial \phi'}{\partial x} - 2w' \right] ,
\]
where \((u', v', w')\) and \((f_x', f_y', f_z')\) are the dimensionless velocity and Lorentz force in the Cartesian coordinate system, respectively.

The dimensionless governing equations of scalar electric potential become
\[
\nabla^2 \phi_x' = \frac{\partial u'}{\partial z} - \frac{\partial w'}{\partial x} ,
\]
\[
\nabla^2 \phi_y' = \frac{\partial w'}{\partial y} - \frac{\partial v'}{\partial z} .
\]
The boundary conditions for the electric potential are determined by \( \mathbf{j} \cdot \mathbf{n} = 0 \), as

\[
\frac{\partial \phi'_e}{\partial n} = \left[ -\varepsilon' \mathbf{e}_x + (\text{Re}_a \text{Pr} \cdot \mathbf{y}' + u') \mathbf{e}_z \right] \mathbf{n}, \quad (32)
\]

\[
\frac{\partial \phi'_e}{\partial n} = \left[ \varepsilon' \mathbf{e}_y + (\text{Re}_a \text{Pr} \cdot \mathbf{x}' - v') \mathbf{e}_z \right] \mathbf{n}. \quad (33)
\]

2.2. Rotating magnetic field \( \Phi_1, \Phi_2 \) model with the skin effect

To investigate the influence of the skin effect in rotating magnetic field, the rotating magnetic field as in equation (17) is applied to the cylindrical liquid bridge. The dimensionless components of the Lorentz force are derived as:

\[
f'_x = \frac{\text{Ta}}{C \cdot \text{Re}_a} \left( \frac{\partial \phi'_e}{\partial x} \hat{B}_{y, \text{lin}} + \frac{\partial \phi'_e}{\partial z} \hat{B}_{z, \text{lin}} \right) + \frac{\text{Ta}}{C^2 \cdot \text{Re}_a} \left[ v' (\hat{B}_{y, \text{lin}} \cdot \hat{B}_{y, \text{lin}} + \hat{B}_{z, \text{lin}} \cdot \hat{B}_{z, \text{lin}}), \quad (34) \right.
\]

\[
\left. -u' (\hat{B}_{z, \text{lin}}^2 + \hat{B}_{y, \text{lin}}^2) \right] \quad - \frac{\text{Ta} \cdot \text{Pr}}{C} (\hat{B}_{x, \text{lin}} \cdot \hat{A}_{x, \text{lin}} + \hat{B}_{y, \text{lin}} \cdot \hat{A}_{y, \text{lin}}) - \frac{\text{Ta} \cdot \text{Pr}}{C} (\hat{B}_{y, \text{lin}} \cdot \hat{A}_{z, \text{lin}} + \hat{B}_{z, \text{lin}} \cdot \hat{A}_{x, \text{lin}}) - \frac{\text{Ta}}{C^2 \cdot \text{Re}_a} \left[ v' (\hat{B}_{x, \text{lin}} \cdot \hat{B}_{x, \text{lin}} + \hat{B}_{y, \text{lin}} \cdot \hat{B}_{z, \text{lin}}), \quad (35) \right.
\]

\[
\left. -u' (\hat{B}_{z, \text{lin}}^2 + \hat{B}_{y, \text{lin}}^2) \right] \quad \left. + \frac{\text{Ta} \cdot \text{Pr}}{C} (\hat{B}_{x, \text{lin}} \cdot \hat{A}_{x, \text{lin}} + \hat{B}_{y, \text{lin}} \cdot \hat{A}_{y, \text{lin}}) + \frac{\text{Ta}}{C^2 \cdot \text{Re}_a} \left[ v' (\hat{B}_{x, \text{lin}} \cdot \hat{B}_{x, \text{lin}} + \hat{B}_{y, \text{lin}} \cdot \hat{B}_{z, \text{lin}}), \quad (36) \right. \right.
\]

Taking \( \mathbf{V} \cdot \mathbf{j} = 0 \), the dimensionless governing equation of scalar electric potential is derived as:

\[
\nabla^2 \phi'_e = \left[ \hat{B}_{x, \text{lin}} \left( \frac{\partial \phi'_e}{\partial x} \mathbf{e}_x + \frac{\partial \phi'_e}{\partial y} \mathbf{e}_y + \frac{\partial \phi'_e}{\partial z} \mathbf{e}_z \right) + w' \hat{A}_{x, \text{lin}} \right]. \quad (37)
\]

\[
\nabla^2 \phi'_e = \left[ \hat{B}_{x, \text{lin}} \left( \frac{\partial \phi'_e}{\partial x} \mathbf{e}_x + \frac{\partial \phi'_e}{\partial y} \mathbf{e}_y + \frac{\partial \phi'_e}{\partial z} \mathbf{e}_z \right) + w' \hat{A}_{x, \text{lin}} \right]. \quad (38)
\]

The boundary conditions for the electric potential are:

\[
\frac{\partial \phi'_e}{\partial n} = \left[ -\frac{1}{C} \varepsilon' \hat{B}_{z, \text{lin}} \mathbf{e}_x + \frac{1}{C} \varepsilon' \hat{B}_{x, \text{lin}} \mathbf{e}_y \right] \mathbf{n}, \quad (39)
\]

In equations (34)-(40), \( C = 2 + \frac{K^2}{96} + \frac{2K^2}{9216} \), and \( \hat{B}_{x, \text{lin}}, \hat{B}_{y, \text{lin}}, \hat{B}_{z, \text{lin}}, \hat{A}_{x, \text{lin}}, \hat{A}_{y, \text{lin}}, \hat{A}_{z, \text{lin}} \) and \( \hat{A}_{x, \text{lin}}, \hat{A}_{y, \text{lin}}, \hat{A}_{z, \text{lin}} \) are exhibited in Appendix B.

The velocity and temperature boundary conditions for two models are

Upper boundary: \( \hat{U} = 0, \hat{T} = 1 \);

Lower boundary: \( \hat{U} = 0, \hat{T} = 0 \);

The circumference boundary: the surface is impervious to flows of mass, momentum and energy.

Initial conditions are: \( \hat{U} = 0, \hat{T} = 0.5, \phi'_1 = 0 \) and \( \phi'_2 = 0 \).

3. Results and Discussion

The dimensionless governing equations (24)-(26) were discretized by the finite volume method. The two-order upwind scheme was applied to the convection term, and the flow field was resolved by the SIMPLE algorithm. The grid convergence was checked as indicated in Table 1, and a non-uniform grid \((60 \times 68 \times 60)\) was adopted. The dimensionless parameters in the present simulations were \( \text{Pr} = 0.01, \text{Re}_a = 2.2 \times 10^4, \text{Ta} = 1.0 \times 10^4 \) and \( K = 0.8 \).

In the cylindrical liquid bridge, convection was driven by the rotating magnetic field and the consequent stirring action of the Lorentz force in the azimuthal direction (which is the same direction of the applied rotating magnetic field), and an axisymmetric rotating flow along the liquid bridge axis was observed, as shown in Fig. 2. The convection in the \( r-z \) plane, as exhibited in Figs. 3(a-c), was a two-vortex-pairs structure induced by centrifugal force and pressure gradient. Under the rotating magnetic field, the skin effect leads to a non-uniform magnetic distribution in the melt, and the magnetic field achieves its maximum strength near the melt surface. Thus, in the

<table>
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<th>Grid</th>
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<th>Maximal Azimuthal Velocity</th>
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</tr>
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</table>
finite length model of the rotating magnetic field, the comparison between the result from the model without the skin effect (i.e. equations (27-29)) and that from the model with the skin effect (i.e. equations (34-36)) indicated an obvious deviation in the azimuthal velocity with increasing dimensionless shielding parameter $K$, as exhibited in Figs. 3(a-c). Also, an obvious deviation in temperature with increasing dimensionless shielding parameter $K$ was observed, as indicated in Fig. 4.

In the flow driven by the rotating magnetic field, the main flow was in the azimuthal direction, and the secondary flow was relatively weak, which facts are exhibited in Figs. 5-6. Because of the skin effect, the magnetic field cannot penetrate into the melt volume without change, and the magnetic lines bend to the cylinder surface more obviously with increasing $K$; therefore, in ignoring the skin effect for a large $K$, an obvious deviation was observed. To quantitatively estimate the error by ignoring the skin effect, the relative error $\zeta$ was defined as $\zeta = \frac{|u_2 - u_1|}{u_1} \times 100\%$, where $u_2$ is the velocity considering the skin effect, and $u_1$ is the velocity without the skin effect. The maximal relative deviation was located near the symmetry axial even though the absolute velocity deviation appeared at the free surface, as shown in Figs. 5(a-c). For $K=2$, the maximal relative velocity error was 6.1 %, and the maximal relative errors of the radial and azimuthal velocities were 15.9 and 7.7 %, respectively. For $K=8$, the maximal relative velocity

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**Fig. 2.** Velocity vector in $z=0.5$ generated by rotating magnetic field.

**Fig. 3.** Velocity vector (left) and azimuthal velocity contours (right) at $x=0$ for rotating magnetic field $\Phi_1-\Phi_2$ model: (a) without skin effect, (b) $K=2$, (c) $K=8$.

**Fig. 4.** (Color online) Temperature distribution at $z=0.7$ for different $K$. 

- **Without skin effect**
- **$K=2$**
- **$K=8$**

---
error was 22.1 %, and the maximal relative errors of the radial and azimuthal velocities are 41.2 and 25.9 %, respectively.

Figs. 6(a–c) illustrate the maximal velocity of melt convection driven by the rotating magnetic field for different K (K = 0 for the case ignoring the skin effect). With the rise of K, the maximal azimuthal velocity increased, while the maximal axial and radial velocities decreased. The maximal azimuthal, axial and radial velocities for K = 8 were 1.077, 0.962 and 0.879 times those for K = 0, respectively. Therefore, the skin effect of the rotating magnetic field cannot be ignored.
4. Conclusion

In terms of Bessel functions, an analytical solution for the distribution of the rotating magnetic field was derived, and the first four terms in the power series expansions of the Kelvin functions were adopted. Then the rotating magnetic field \( \Phi \) model with the skin effect was proposed. The flow driven by the rotating magnetic field was simulated numerically, and the influence of the skin effect was investigated. The results indicate that an obvious deviation in velocity appears when the skin effect is ignored. The rotating magnetic field model ignoring the skin effect leads to a smaller azimuthal velocity and larger radial and axial velocity components.

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Appendix A

From the equation (15) and (16), the distribution of rotating magnetic field with the skin effect in melt are

\[
B_r = \text{Re}(\frac{R R_0}{r} \text{ber}_1(\sqrt{K} r) + i \text{bei}_1(\sqrt{K} r) e^{i(\omega t - \theta)})
\]

\[
B_0 = \text{Re}(\frac{R R_0}{R} \text{ber}_1(\sqrt{K} R) + i \text{bei}_1(\sqrt{K} R))
\]

In this study, the first four terms in the power series expansions of the Kelvin functions are reserved,

\[
\text{ber}_1(\sqrt{K} r) = -\frac{1}{\sqrt{2}} \left( \frac{\sqrt{K} r}{2 R} \right)^3 \left( \frac{\sqrt{K} r}{2 R} \right)^2 \left( \frac{\sqrt{K} r}{2 R} \right)
\]

\[
\text{bei}_1(\sqrt{K} r) = -\frac{1}{\sqrt{2}} \left( \frac{\sqrt{K} r}{2 R} \right)^3 \left( \frac{\sqrt{K} r}{2 R} \right)^2 \left( \frac{\sqrt{K} r}{2 R} \right)
\]

\[
\text{ber}_1(\sqrt{K} R) = -\frac{1}{\sqrt{2}} \left( \frac{\sqrt{K} R}{2 R} \right)^3 \left( \frac{\sqrt{K} R}{2 R} \right)^2 \left( \frac{\sqrt{K} R}{2 R} \right)
\]

\[
\text{bei}_1(\sqrt{K} R) = -\frac{1}{\sqrt{2}} \left( \frac{\sqrt{K} R}{2 R} \right)^3 \left( \frac{\sqrt{K} R}{2 R} \right)^2 \left( \frac{\sqrt{K} R}{2 R} \right)
\]

Hence,

\[
B_r = \frac{B_0}{2 + K^2 + 2K^6}
\]

\[
B_0 = \text{Re}(\frac{R R_0}{R} \text{ber}_1(\sqrt{K} R) + i \text{bei}_1(\sqrt{K} R))
\]

The magnetic field with the skin effect in the melt in the Cartesian coordinate system is transformed from

\[
B_x = B_r \cos \theta - B_\theta \sin \theta, B_y = B_r \sin \theta + B_\theta \cos \theta, x = r \cos \theta \text{ and } y = r \sin \theta \text{ hence}
\]

\[
\begin{bmatrix}
\cos(\omega t - \theta) \\
\sin(\omega t - \theta)
\end{bmatrix}
\]

\[
\frac{K}{4} \left[ \frac{K^2}{4608} + \frac{1}{R^2} \left( \frac{K^3}{678} + \frac{K^3}{32} \right) \right]
\]

\[
\frac{K}{4} \left[ \frac{K^2}{4608} + \frac{1}{R^2} \left( \frac{K^3}{678} + \frac{K^3}{32} \right) \right]
\]

\[
\frac{K}{4} \left[ \frac{K^2}{4608} + \frac{1}{R^2} \left( \frac{K^3}{678} + \frac{K^3}{32} \right) \right]
\]
The magnetic field with the skin effect in the melt is denoted as

$$\mathbf{B}_{\text{rot}}(x, y, t) = \frac{B_t}{2 + \frac{K^2}{96} + \frac{2K^6}{9216}} \left[ (\sin \omega t - B_{\text{cos}} + \cos \omega t - B_{\text{sin}})e_x \right]$$

where $B_{\text{sin}}, B_{\text{cos}}, B_{y \text{sin}}$ and $B_{y \text{cos}}$ are described as

$$B_{\text{sin}} = \frac{K^3}{4} + \frac{1}{R^2} \left( \frac{K^3}{4608} + \frac{1}{R^2} \right) \left( \frac{K^3}{768} - \frac{K^3}{4} \right) (x^2 + 3y^2)$$

$$+ \frac{1}{R^2} \left( \frac{K^3}{884736} - \frac{K^3}{768} \right) (x^2 + y^2)^2 (x^2 + 5y^2)$$

$$+ \frac{1}{R^2} \left( \frac{K^3}{4608} - \frac{K^3}{884736} \right) (x^2 + y^2)^2 (x^2 + 7y^2)$$

$$- \frac{1}{R^2} \left( -\frac{12K^6}{9216} - \frac{K^4}{6144} \right) (x^2 + y^2)^2 xy$$

$$B_{\text{cos}} = 2 - \frac{K^2}{96} + \frac{1}{R^2} \left( \frac{K^2}{12} - \frac{K^2}{36864} \right) (x^2 + 3y^2)$$

$$+ \frac{1}{R^2} \left( \frac{2K^4}{192^2} - \frac{K^4}{96} \right) (x^2 + y^2)^2 (x^2 + 5y^2)$$

$$+ \frac{1}{R^2} \left( \frac{2K^6}{9216} - \frac{K^6}{36864} \right) (x^2 + y^2)^2 (x^2 + 7y^2)$$

$$+ \frac{1}{R^2} \left( \frac{K^2}{K^3} + \frac{1}{R^2} \right) \left( \frac{K^2}{221184} - \frac{K^2}{192} \right) (x^2 + y^2) xy$$

$$+ \frac{1}{R^2} \left( \frac{K^3}{768} - \frac{K^3}{147456} \right) (x^2 + y^2)^2 xy$$

The following functions are functions of dimensionless $x^*, y^*$, and the symbol of "**" is omitted for the sake of convenience.

In calculating the dimensionless Lorentz force by the rotating magnetic field with the skin effect $\mathbf{B}_{\text{rot}}(x, y, t)$, we have

$$\mathbf{B}_{\text{rot}} = \frac{K^3}{4} + \frac{1}{R^2} \left( \frac{K^3}{4608} + \frac{1}{R^2} \right) \left( \frac{K^3}{768} - \frac{K^3}{4} \right) (x^2 + 3y^2)$$

$$+ \frac{1}{R^2} \left( \frac{K^3}{884736} - \frac{K^3}{768} \right) (x^2 + y^2)^2 (x^2 + 5y^2)$$

$$+ \frac{1}{R^2} \left( \frac{K^3}{4608} - \frac{K^3}{884736} \right) (x^2 + y^2)^2 (x^2 + 7y^2)$$

$$- \frac{1}{R^2} \left( -\frac{12K^6}{9216} - \frac{K^4}{6144} \right) (x^2 + y^2)^2 xy$$

$$\mathbf{B}_{\text{rot}} = 2 - \frac{K^2}{96} + \frac{1}{R^2} \left( \frac{K^2}{12} - \frac{K^2}{36864} \right) (x^2 + 3y^2)$$

$$+ \frac{1}{R^2} \left( \frac{2K^4}{192^2} - \frac{K^4}{96} \right) (x^2 + y^2)^2 (x^2 + 5y^2)$$

$$+ \frac{1}{R^2} \left( \frac{2K^6}{9216} - \frac{K^6}{36864} \right) (x^2 + y^2)^2 (x^2 + 7y^2)$$

$$+ \frac{1}{R^2} \left( \frac{K^2}{K^3} + \frac{1}{R^2} \right) \left( \frac{K^2}{221184} - \frac{K^2}{192} \right) (x^2 + y^2) xy$$

$$+ \frac{1}{R^2} \left( \frac{K^3}{768} - \frac{K^3}{147456} \right) (x^2 + y^2)^2 xy$$
\[
\hat{B}_{y \text{sin}} = -2 \frac{K^2}{96} + \left( \frac{K^3}{32} - \frac{K^4}{36864} \right) (3x^2 + y^2)
\]
\[
+ \left( \frac{2K^4}{192} - \frac{K^2}{96} \right) (x^2 + y^2) (5x^2 + y^2)
\]
\[
+ \left( \frac{2K^6}{9216} - \frac{K^4}{36864} \right) (x^2 + y^2)^2 (7x^2 + y^2)
\]
\[
- \left( \frac{K^3}{384} - \frac{K^3}{221184} \right) (x^2 + y^2)(x^2 + y^2)xy
\]
\[
- \left( \frac{K^3}{768} - \frac{K^3}{147456} \right) (x^2 + y^2)^2 xy
\]
\[
\hat{B}_{y \text{cos}} = -\left( \frac{K^3}{4} - \frac{K^3}{4608} \right) - \left( \frac{K^3}{36864} - \frac{K^3}{768} \right) (x^2 + y^2)
\]
\[
- \left( \frac{K^3}{884376} - \frac{K^3}{768} \right) (x^2 + y^2)^2 (7x^2 + y^2)
\]
\[
- \left( \frac{K^4}{4608} - \frac{K^4}{884376} \right) (x^2 + y^2)^2 (7x^2 + y^2)
\]
\[
- \left( \frac{K^4}{192} - \frac{K^4}{24} \right) y (x^2 + y^2)(x^2 + y^2)xy
\]
\[
- \left( \frac{12K^6}{9216} - \frac{K^4}{6144} \right) (x^2 + y^2)^2 xy
\]
\[
\phi_{x \text{in}} = \left( \frac{K^3}{96} - \frac{K^3}{884376} \right) x - \left( \frac{K^3}{36864} - \frac{K^3}{768} \right) y
\]
\[
+ \left( \frac{K^3}{32} - \frac{K^3}{18432} \right) (x^2 + y^2) y + \left( \frac{K^4}{96} - \frac{K^4}{884376} \right) (x^2 + y^2) y
\]
\[
- \left( \frac{K^4}{192} - \frac{K^4}{24} \right) (x^2 + y^2) y - \left( \frac{K^4}{36864} - \frac{K^4}{768} \right) (x^2 + y^2) y
\]
\[
\phi_{x \text{cos}} = \left( \frac{K^3}{96} - \frac{K^3}{4608} \right) x + \left( \frac{K^3}{36864} - \frac{K^3}{768} \right) y
\]
\[
+ \left( \frac{K^3}{32} - \frac{K^3}{18432} \right) (x^2 + y^2) y + \left( \frac{K^4}{96} - \frac{K^4}{884376} \right) (x^2 + y^2) y
\]
\[
+ \left( \frac{K^4}{192} - \frac{K^4}{24} \right) (x^2 + y^2) y + \left( \frac{K^4}{36864} - \frac{K^4}{768} \right) (x^2 + y^2) y
\]
\[
A_{x \text{sin}} = \left( \frac{K^3}{4} - \frac{K^3}{4608} \right) x + \left( \frac{K^3}{32} - \frac{K^3}{36864} \right) (x^2 + y^2) y
\]
\[
+ \left( \frac{2K^4}{192} - \frac{K^2}{96} \right) (x^2 + y^2) y + \left( \frac{K^4}{12} - \frac{K^4}{36864} \right) (x^2 + y^2) y
\]
\[
+ \left( \frac{K^3}{768} - \frac{K^3}{884376} \right) (x^2 + y^2) y + \left( \frac{K^3}{36864} - \frac{K^3}{768} \right) (x^2 + y^2) y
\]
\[
A_{x \text{cos}} = \left( 2 - \frac{K^2}{96} \right) x - \left( \frac{K^3}{4} - \frac{K^3}{4608} \right) y
\]
\( f_x' \): Dimensionless component of the Lorentz force in the \( z \) direction

References